

Canned WET goods:  
Low-energy BSM constraints for model builders

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Heavy flavour aspects – 24/04/2023

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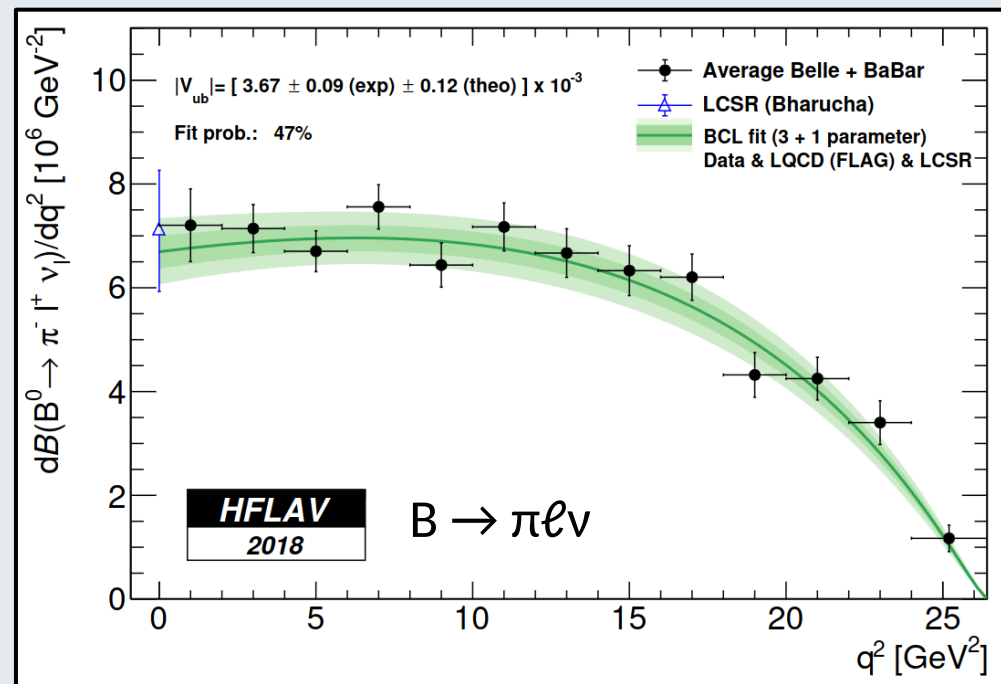
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Based on: Leljak, Melić, Novak, MR, van Dyk 2302.05268

- Low-energy analyses deal with  **$O(100)$  hadronic nuisance parameters**
    - QCD is non-perturbative
    - Lattice calculations are complicated
    - Other QCD methods (LCSR, QM) have uncontrolled uncertainties
- This is not something model builders want/should deal with
- **Question:** How can low-energy information be passed on to model builders with minimum dilution and modification?
- I will try to provide a partial answer in the context of  $b \rightarrow u\ell\nu$

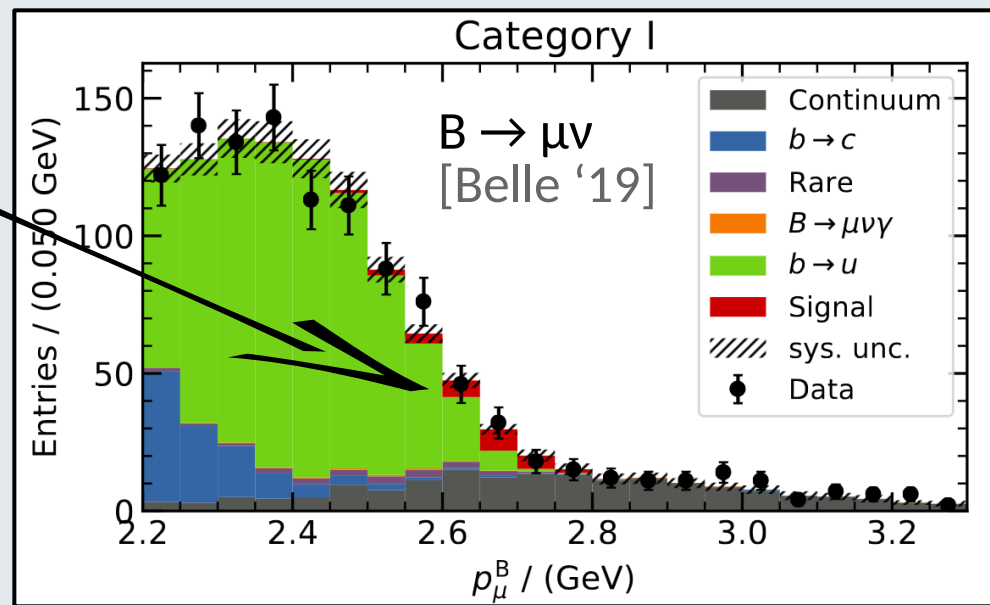
# Experimental inputs for $b \rightarrow u\ell\nu$ (I)

- $B \rightarrow \pi\ell\nu$ : HFLAV '19 average of BaBar '10 and '12 and Belle '10 and '13 measurements  $\rightarrow$  13  $q^2$  bins
- $B \rightarrow \rho\ell\nu$ : average of BaBar '10 and Belle '13 measurements performed by [Bernlochner et al '21]  $\rightarrow$  11  $q^2$  bins
- $B \rightarrow \omega\ell\nu$ : average of BaBar '12 and Belle '13 measurements performed by [Bernlochner et al '21]  $\rightarrow$  5  $q^2$  bins



# Experimental inputs for $b \rightarrow u\ell\nu$ (II)

- We don't add the following data (yet):
  - $B \rightarrow \mu\nu$ : only observed with  $2.8\sigma$  significance
  - $B \rightarrow \eta^{(\prime)}\ell\nu$ : little statistics and poorly known form factors
  - Inclusive  $b \rightarrow u\ell\nu$ : assume SM in the analysis (more details later)
  - $\Lambda_b \rightarrow p\ell\nu$  and  $B_s \rightarrow K^{(*)}\ell\nu$  are measured by LHCb, but normalized to a  $b \rightarrow c\ell\nu$  mode!



- Assume left-handed neutrinos only

$$\mathcal{H}^{ubl\nu} = -\frac{4G_F}{\sqrt{2}} \tilde{V}_{ub} \sum_i \mathcal{C}_i^\ell \mathcal{O}_i^\ell + \dots + \text{h.c.}$$

- $B \rightarrow \pi$  (pseudoscalar) and  $B \rightarrow \rho$ ,  $B \rightarrow \omega$  (vectors) are sensitive to different combinations of the operators

$$\begin{aligned} \mathcal{O}_{V,L}^\ell &= [\bar{u}\gamma^\mu P_L b] [\bar{\ell}\gamma_\mu P_L \nu], & \mathcal{O}_{V,R}^\ell &= [\bar{u}\gamma^\mu P_R b] [\bar{\ell}\gamma_\mu P_L \nu], \\ \mathcal{O}_{S,L}^\ell &= [\bar{u}P_L b] [\bar{\ell}P_L \nu], & \mathcal{O}_{S,R}^\ell &= [\bar{u}P_R b] [\bar{\ell}P_L \nu], \\ \mathcal{O}_T^\ell &= [\bar{u}\sigma^{\mu\nu} b] [\bar{\ell}\sigma_{\mu\nu} P_L \nu]. \end{aligned}$$

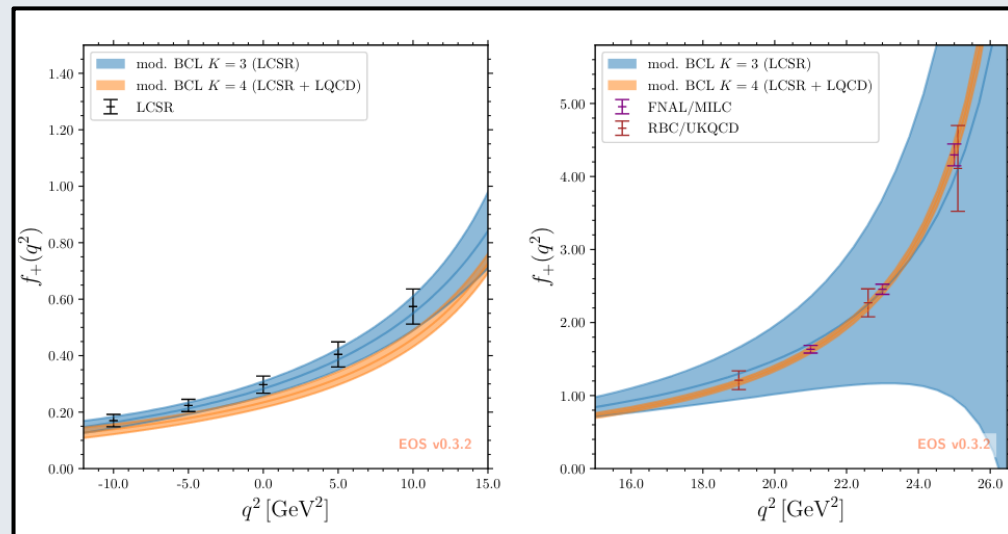
- Only one relevant operator in the SM, normalized to  $V_{ub}$

$$\mathcal{C}_{V,L}^\ell = 1 + \frac{\alpha_e}{\pi} \ln \left( \frac{M_Z}{\mu} \right)$$

# Hadronic inputs

- All the non-perturbativity of QCD appears in the *hadronic form factors*. These functions are known from LQCD and/or LCSR calculations:
  - $B \rightarrow \pi$ , we used the same inputs as [Leljak, Melić, van Dyk '21]
    - LQCD [FNAL+MILC '15] [UKQCD '15] and LCSR [Leljak, Melić, van Dyk '21]
    - BCL parametrization  
→ 12 parameters
  - $B \rightarrow \rho$  and  $B \rightarrow \omega$ 
    - LCSR [Bharucha, Straub, Zwicky '15]
    - BSZ parametrization  
→ 2\*19 parameters

→ 50 nuisance parameters



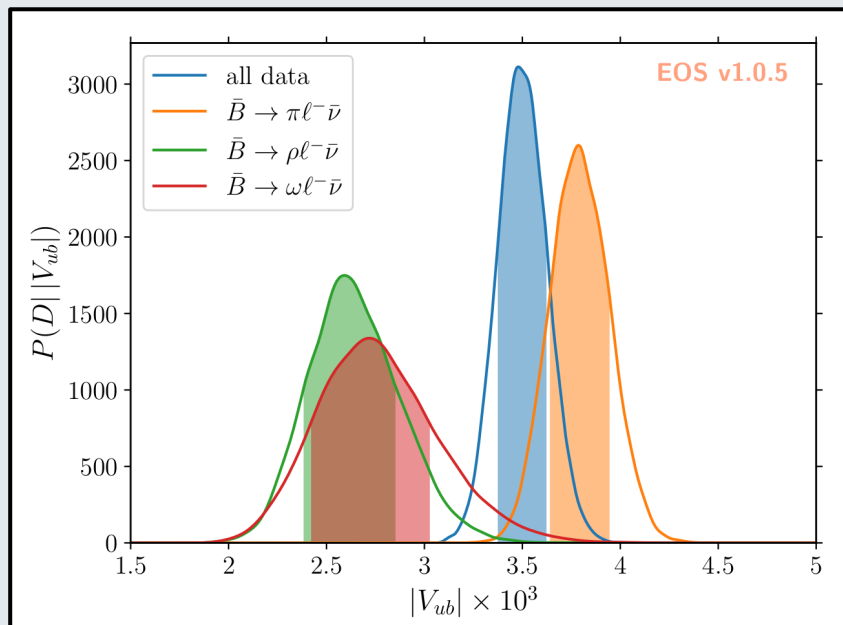
[Leljak, Melić, van Dyk '21]

- Our goals:
  - 1) Determine the consistency of exclusive data and the quality of a  $|V_{ub}|$  extraction
  - 2) Determine whether a BSM explanation of the data is favored over the SM
  - 3) Provide the posterior likelihood of the WCs  $C_i$
- We perform 3 Bayesian analyses:
  - **SM** → only float hadronic inputs: *null hypothesis*
  - **CKM** → float hadronic inputs and extract  $|V_{ub}|$
  - **WET** → float hadronic inputs and extract the WCs  $C_i$
- We sampled using *nested sampling* and **EOS**: [eos.github.io](https://eos.github.io)



# Results: Extraction of $|V_{ub}|$

- Good fits: p values > 52%
- Form factors uncertainties propagate to  $|V_{ub}|$



Data set	Goodness of fit			$ V_{ub}  \times 10^3$
	$\chi^2$	d.o.f.	p value [%]	
$\bar{B} \rightarrow \pi \ell \nu$	27.83	31	62.98	$3.79^{+0.15}_{-0.15}$
$\bar{B} \rightarrow \rho \ell \nu$	5.08	10	88.60	$2.63^{+0.25}_{-0.22}$
$\bar{B} \rightarrow \omega \ell \nu$	3.19	4	52.66	$2.74^{+0.33}_{-0.28}$
all data	52.31	47	27.53	$3.50^{+0.13}_{-0.12}$

- State-of-the-art determinations:
  - Inclusive [HFLAV, PDG, ... '22]  
 $|V_{ub}| = 4.13(12)(13)(18) 10^{-3}$
  - Exclusive [HFLAV, PDG, ... '20]  
 $|V_{ub}| = 3.70(10)(12) 10^{-3}$



# Results: Predictions for $B \rightarrow \ell \nu$

- We can post-dict values for all relevant  $b \rightarrow u \ell \nu$  observable, e.g.

$$\mathcal{B}(\bar{B}^- \rightarrow \tau^- \bar{\nu}) = \left( 7.87_{-0.54}^{+0.58} \Big|_{|V_{ub}|} \pm 0.12 \Big|_{f_B} \right) \times 10^{-5},$$

$$\mathcal{B}(\bar{B}^- \rightarrow \mu^- \bar{\nu}) = \left( 3.54_{-0.24}^{+0.26} \Big|_{|V_{ub}|} \pm 0.05 \Big|_{f_B} \right) \times 10^{-7},$$

$$\mathcal{B}(\bar{B}^- \rightarrow e^- \bar{\nu}) = \left( 8.28_{-0.56}^{+0.61} \Big|_{|V_{ub}|} \pm 0.12 \Big|_{f_B} \right) \times 10^{-12}.$$

- To be compared with the experimental result:

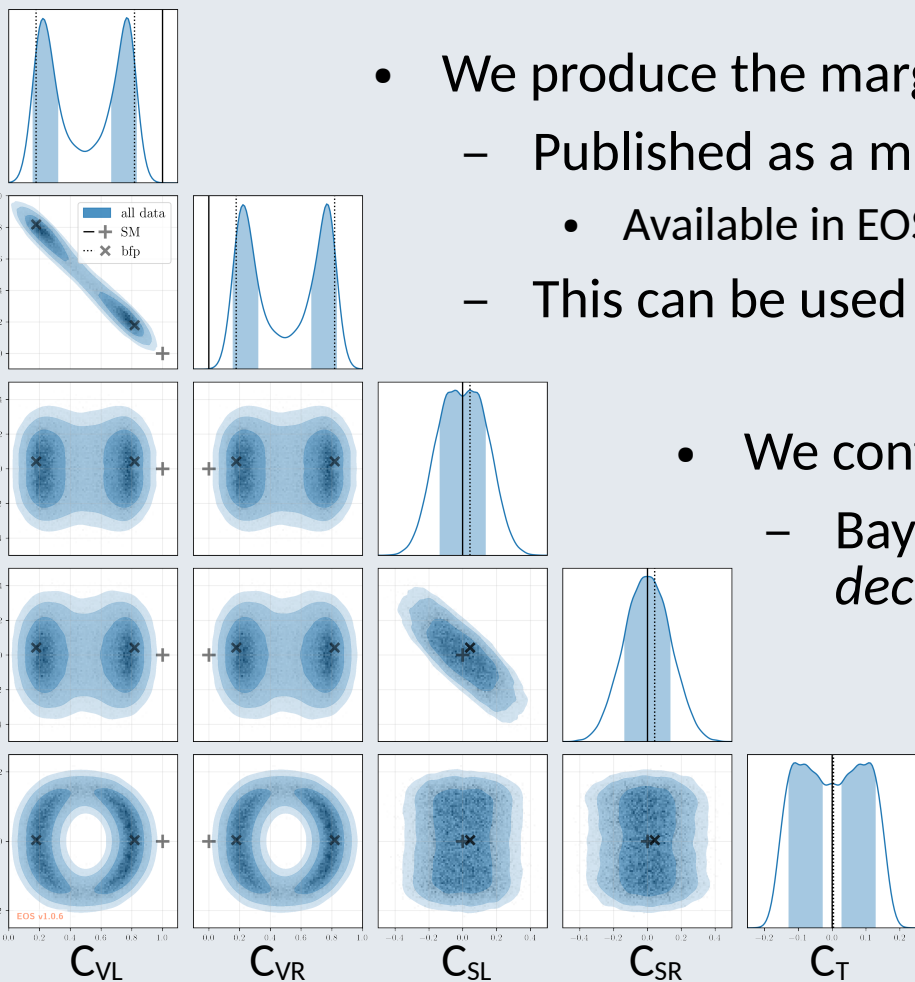
$$\mathcal{B}(\bar{B}^- \rightarrow \mu^- \bar{\nu}) \Big|_{\text{Belle '19}} = (5.3 \pm 2.2) \times 10^{-7}$$

# Results: Global likelihood

- We produce the marginalized posterior distribution of the  $W$ cs
  - Published as a mixture density of multivariate Gaussians
    - Available in EOS as `ub1nu::P(WET)@LMNRvD:2023A`
  - This can be used by model builders to study BSM scenarios

- We confirm the (known) tension in the  $C_{VL} - C_{VR}$  plane
  - Bayesian model comparison finds the **WET** model *decisively in favour* over the **SM** and **CKM** ones

- Future (angular) analyses will help breaking the degeneracy in the  $C_{VL} - C_{VR}$  plane



# Conclusion (I)

- Yes, the  $|V_{ub}|$  inclusive vs. exclusive tension is still present, albeit diminished
- Yes, NP contributions consistently improve the  $b \rightarrow u\ell\nu$  fit
- But more important (in my opinion): this analysis is meant as a **benchmark for future work**:

Testing BSM models cannot be done with  $O(100)$  hadronic nuisance parameters

→ The theory community will need such (up-to-date) WET likelihoods

→ The theory **and** experimental communities will need to agree on

- (1) an exchange format for non Gaussian likelihoods
- (2) hadronic inputs
- (3) observables of interest...

Examples of problems we need to solve/avoid:

- $b \rightarrow u\ell\nu$  inclusive (and exclusive, to a smaller extent) analyses need to **assume SM** for MC production, efficiency calculation...
  - This makes a WET analysis *impossible at present*
  - One possible solution is to reweight MC samples with BSM weights (Hammer, EOS...)
  - But all the analysis steps have to be adapted
- Some experiments suggest an unbinned fit of the WET WC
  - This should only be an **additional piece of information**
    - Global analyses would require the full posterior, including all nuisance parameters which is very hard if not impossible to achieve.